

Can neutral correlations be explained in the Standard Model?

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ASSUME:

- Energies are measured correctly
- Primary particles are neutral, their fraction is

$$1.5\% < \eta < 3.5\%$$

- Sources are at $\gtrsim 400$ Mpc

SHOW:

NO EXPLANATION EXISTS WITHIN THE STANDARD MODEL

To cross the Galactic magnetic field particle lifetime must be

$$\tau > 10 \text{ s} \left(\frac{m}{1 \text{ GeV}} \right) \left(\frac{10^{19} \text{ eV}}{E} \right)$$

SM candidates:

- neutrino
- photon
- neutron
- atoms

Photon and neutrino

- * Photons of energy 10^{19} eV do not reach Earth from 400 Mpc
- * Neutrino is not interacting strongly enough to be the primary particle

Neutron

Weak reactions

On neutrino background

$$\sigma \propto G_F^2 E^2$$

rate

$$R \sim 3 \times 10^{-12} \frac{1}{\text{Mpc}}$$

⇒ out of question: too small

Nuclei photodisintegration

We need the reaction

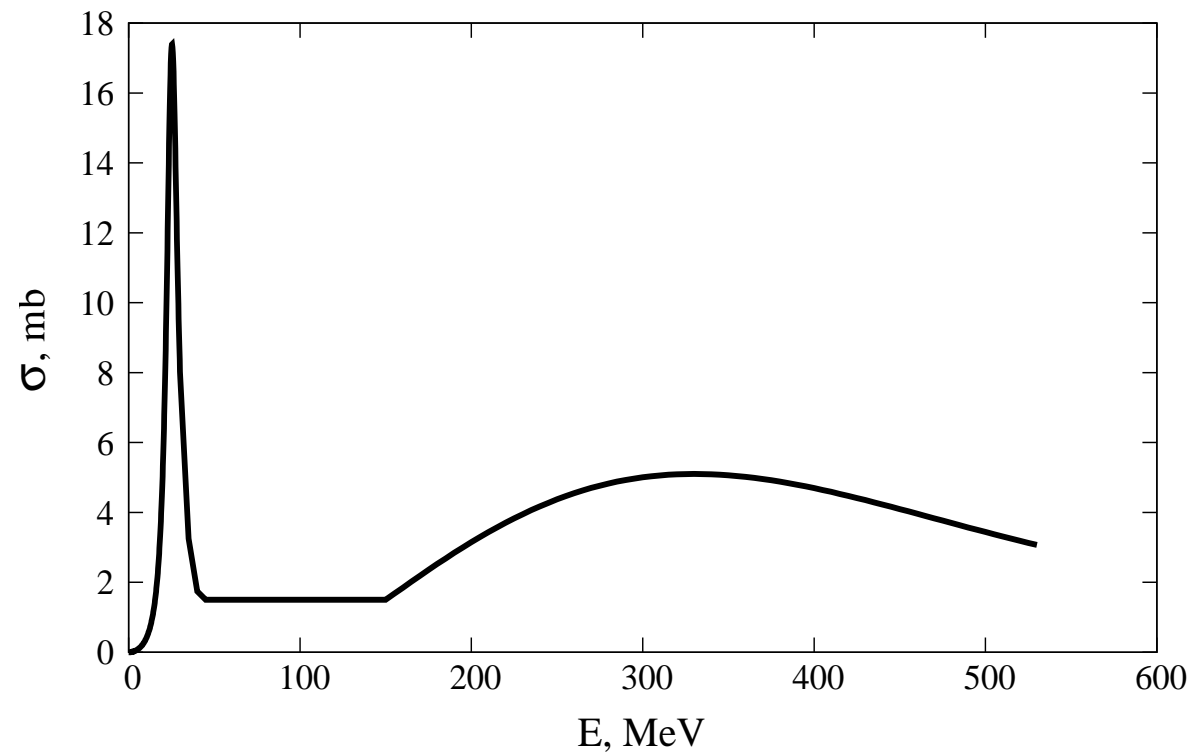


to occur within 100 kpc from Earth

$$R = \int d^3p n(\mathbf{p})(1 - v \cos \theta) \sigma_{\text{NRF}}[\gamma p(1 - v \cos \theta)]$$

$$\sim \frac{2\pi}{\gamma^2} \int_0^\infty dp n(\mathbf{p}) \int_0^{2\gamma p} d\omega' \omega' \sigma_{\text{NRF}}(\omega')$$

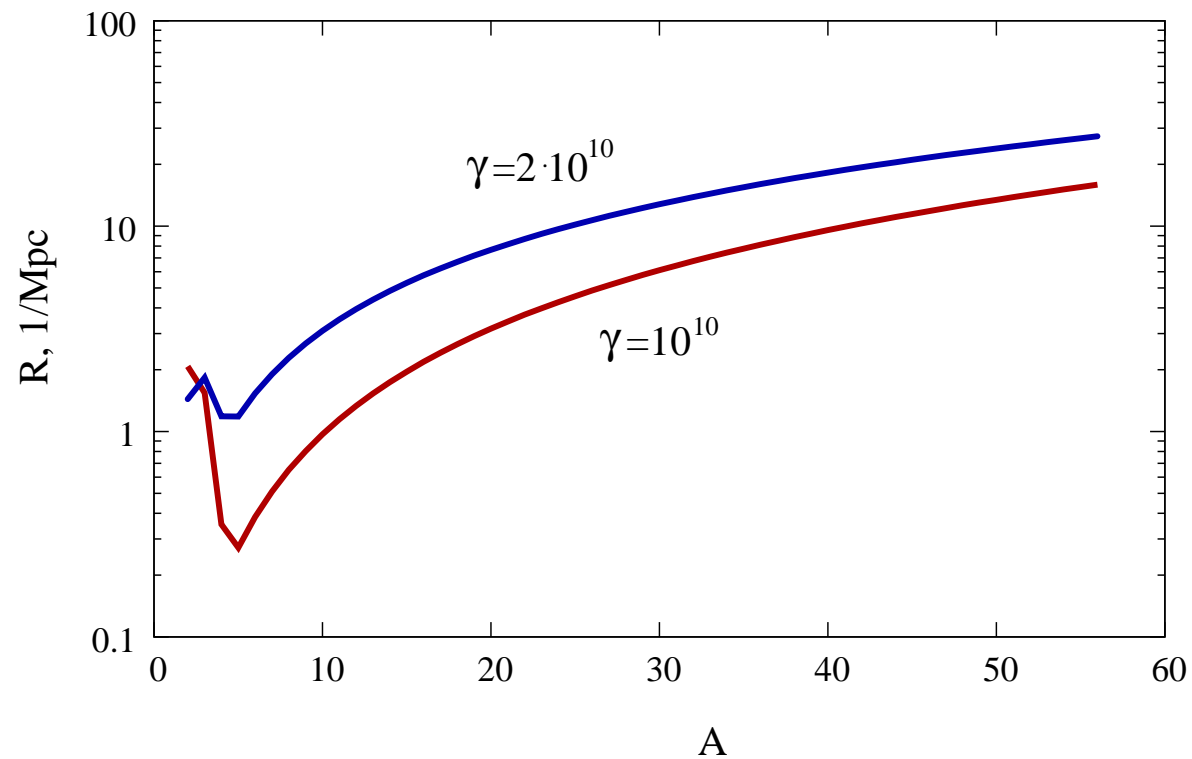
The cross section of photodisintegration is measured, and an analytic parameterization exists.



* Take

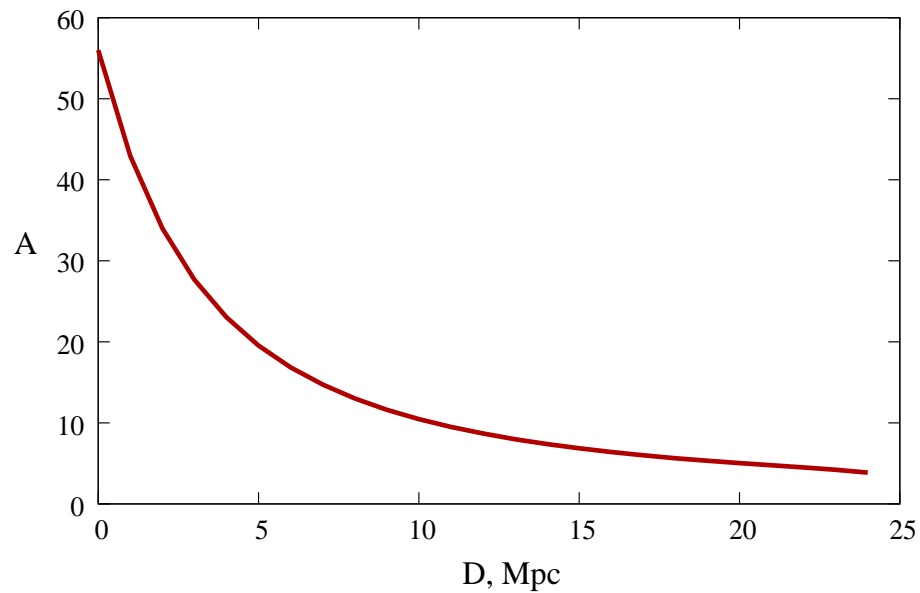
$$\gamma = (1 - 2) \times 10^{10}$$

* Take **only** CMB photons (IR background can only make the rate larger)



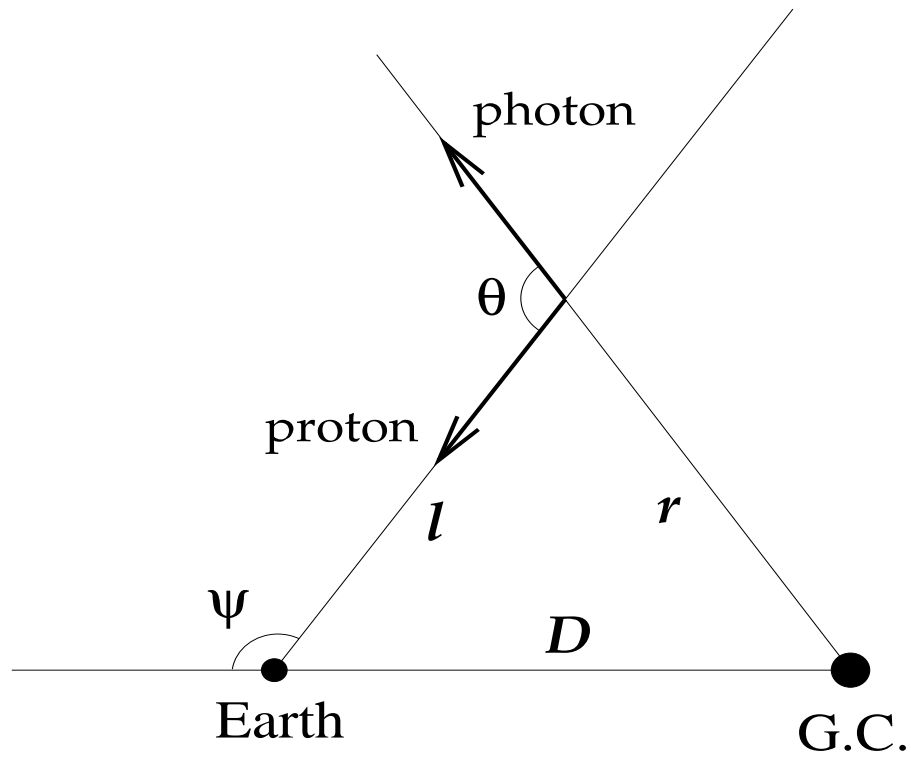
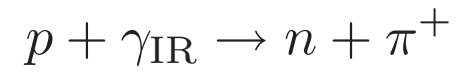
How does A evolve with the distance from the source?

$$\frac{dA}{dt} = -R(A)$$



⇒ No hope to produce neutrons this way

Neutrons by pion photoproduction on Galactic IR



Galactic IR background

Modeled by the point source at the center

$$n(\mathbf{p}) = \frac{I_0}{r^2} \frac{p^{1.65}}{e^{-p/T} - 1} \delta(\mathbf{n} - \mathbf{n}_0)$$

IR luminosity within solar orbit is measured,

$$L_G = 2 \times 10^{10} L_\odot = 8 \times 10^{36} \text{W}$$

$$\Rightarrow I_0 = \frac{L_G}{60.3 T^{5.65}}; \quad T = 23.3^\circ\text{K}$$

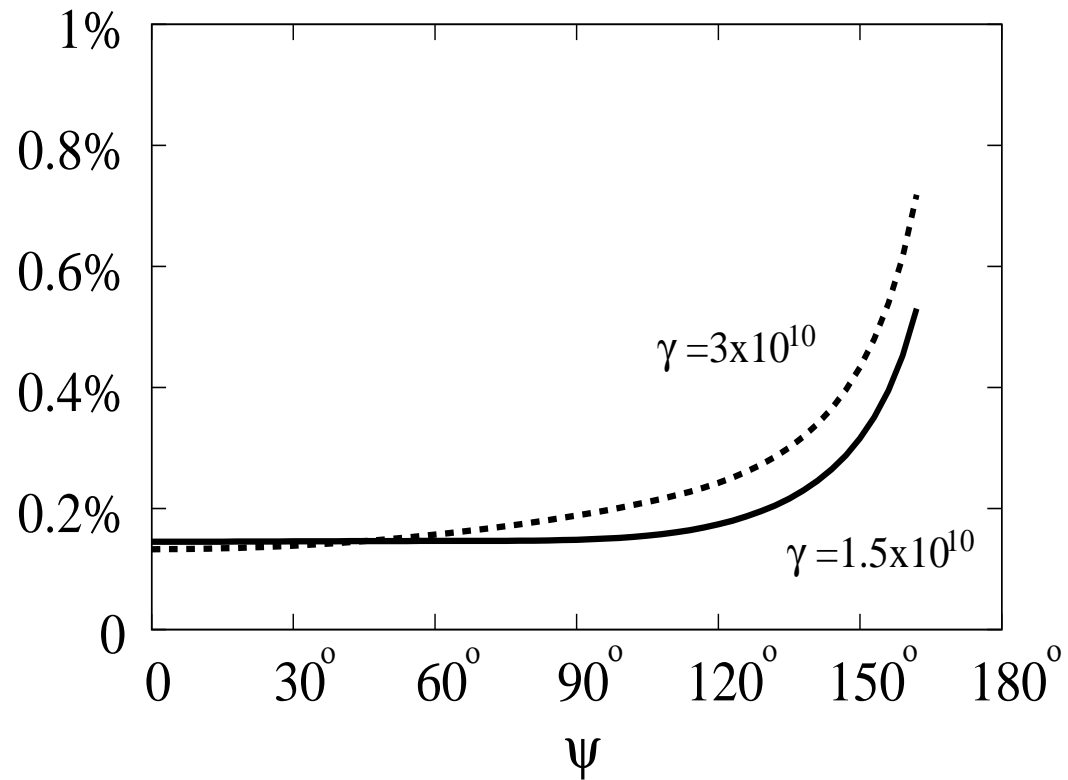
$$P = \frac{1}{60.3} \frac{L_G}{T} \int_0^\infty dl \frac{e^{-l/\lambda}}{r^2} (1 - v \cos \theta) \langle \sigma \rangle (T_{\text{eff}}),$$

where

$$\langle \sigma \rangle (T) = \frac{1}{T^{4.65}} \int \frac{\epsilon^{3.65} \sigma(\epsilon) d\epsilon}{e^{\epsilon/T} - 1}$$

and

$$T_{\text{eff}}(\theta) = \gamma(1 - v \cos \theta)T$$



⇒ An order (only!) of magnitude too small

Atoms

Rate of formation by creating the e^+e^- -pair and dressing

$$R_{\text{formation}} \sim 10^{-5} \frac{1}{\text{Mpc}}.$$

Rate of decay on CMB:

$$R_{\text{decay}} \sim 100 \frac{1}{\text{Mpc}}.$$

\implies Fraction of atoms is too small

CONCLUSIONS

- * There seem to be no explanation within the SM
- * The neutron production on the Galactic IR background came closest – only about an order of magnitude too small.